Part 2: Rigorous derivation of ITG growth rate & threshold (in a simple limit) starting from the Gyrokinetic Eq.

Notes by Greg Hammett, 2007 Presented at the 2007 UCLA Winter School of the Center for Multiscale Plasma Dynamics (CPMD). Our starting point will be the electrostatic Gyrokinetic Eq. written in a Drift-Kinetic-like form for the full, gyro-averaged, gooding center density $f(R, v_1, p, t)$:

$$\frac{\partial \widetilde{f}}{\partial t} + (v_{||}\hat{\mathbf{b}} + \mathbf{v}_{E} + \mathbf{v}_{d}) \cdot \nabla \widetilde{f} + \left(\frac{q}{m}E_{||} - \mu \nabla_{||}B + v_{||}(\hat{\mathbf{b}} \cdot \nabla \hat{\mathbf{b}}) \cdot \mathbf{v}_{E}\right) \frac{\partial \widetilde{f}}{\partial v_{||}} = 0$$

$$\mathbf{v}_{d} = \frac{\mathbf{v}_{\parallel}^{2}}{\Omega}\hat{\mathbf{b}} \times (\hat{\mathbf{b}} \cdot \nabla \hat{\mathbf{b}}) + \frac{\mu}{\Omega}\hat{\mathbf{b}} \times \nabla B \approx \frac{\mathbf{v}_{\parallel}^{2} + \mathbf{v}_{\perp}^{2}/2}{\Omega B^{2}} B \times \nabla B$$

$$\mathbf{v}_{d} = \frac{\mathbf{v}_{\parallel}^{2}}{\Omega}\hat{\mathbf{b}} \times (\hat{\mathbf{b}} \cdot \nabla \hat{\mathbf{b}}) + \frac{\mu}{\Omega}\hat{\mathbf{b}} \times \nabla B \approx \frac{\mathbf{v}_{\parallel}^{2} + \mathbf{v}_{\perp}^{2}/2}{\Omega B^{2}} B \times \nabla B$$

$$\mathbf{v}_{d} = \frac{\mathbf{v}_{\parallel}^{2}}{\Omega}\hat{\mathbf{b}} \times (\hat{\mathbf{b}} \cdot \nabla \hat{\mathbf{b}}) + \frac{\mu}{\Omega}\hat{\mathbf{b}} \times \nabla B \approx \frac{\mathbf{v}_{\parallel}^{2} + \mathbf{v}_{\perp}^{2}/2}{\Omega B^{2}} B \times \nabla B$$

$$\mathbf{v}_{d} = \frac{\mathbf{v}_{\parallel}^{2}}{\Omega}\hat{\mathbf{b}} \times (\hat{\mathbf{b}} \cdot \nabla \hat{\mathbf{b}}) + \frac{\mu}{\Omega}\hat{\mathbf{b}} \times \nabla B \approx \frac{\mathbf{v}_{\parallel}^{2} + \mathbf{v}_{\perp}^{2}/2}{\Omega B^{2}} B \times \nabla B$$

$$\mathbf{v}_{d} = \frac{\mathbf{v}_{\parallel}^{2}}{\Omega}\hat{\mathbf{b}} \times (\hat{\mathbf{b}} \cdot \nabla \hat{\mathbf{b}}) + \frac{\mu}{\Omega}\hat{\mathbf{b}} \times \nabla B \approx \frac{\mathbf{v}_{\parallel}^{2} + \mathbf{v}_{\perp}^{2}/2}{\Omega B^{2}} B \times \nabla B$$

$$\mathbf{v}_{d} = \frac{\mathbf{v}_{\parallel}^{2}}{\Omega}\hat{\mathbf{b}} \times (\hat{\mathbf{b}} \cdot \nabla \hat{\mathbf{b}}) + \frac{\mu}{\Omega}\hat{\mathbf{b}} \times \nabla B \approx \frac{\mathbf{v}_{\parallel}^{2} + \mathbf{v}_{\perp}^{2}/2}{\Omega B^{2}} B \times \nabla B$$

$$\mathbf{v}_{d} = \frac{\mathbf{v}_{\parallel}^{2}}{\Omega}\hat{\mathbf{b}} \times (\hat{\mathbf{b}} \cdot \nabla \hat{\mathbf{b}}) + \frac{\mu}{\Omega}\hat{\mathbf{b}} \times \nabla B \approx \frac{\mathbf{v}_{\parallel}^{2} + \mathbf{v}_{\perp}^{2}/2}{\Omega B^{2}} B \times \nabla B$$

$$\mathbf{v}_{d} = \frac{\mathbf{v}_{\parallel}^{2}}{\Omega}\hat{\mathbf{b}} \times (\hat{\mathbf{b}} \cdot \nabla \hat{\mathbf{b}}) + \frac{\mu}{\Omega}\hat{\mathbf{b}} \times \nabla B \approx \frac{\mathbf{v}_{\parallel}^{2} + \mathbf{v}_{\perp}^{2}/2}{\Omega B^{2}} B \times \nabla B$$

$$\mathbf{v}_{d} = \frac{\mathbf{v}_{\parallel}^{2}}{\Omega}\hat{\mathbf{b}} \times (\hat{\mathbf{b}} \cdot \nabla \hat{\mathbf{b}}) + \frac{\mu}{\Omega}\hat{\mathbf{b}} \times \nabla B \times \nabla B \times \nabla B$$

$$\mathbf{v}_{d} = \frac{\mathbf{v}_{\parallel}^{2}}{\Omega B^{2}} + \frac{\mathbf{v}_{\parallel}^{2}}{\Omega B^{2}} + \frac{\mathbf{v}_{\parallel}^{2}}{\Omega B^{2}} B \times \nabla B$$

$$\mathbf{v}_{d} = \frac{\mathbf{v}_{\parallel}^{2}}{\Omega}\hat{\mathbf{b}} \times (\hat{\mathbf{b}} \cdot \nabla \hat{\mathbf{b}}) + \frac{\mu}{\Omega}\hat{\mathbf{b}} \times \nabla B \times \nabla B \times \nabla B \times \nabla B$$

$$\mathbf{v}_{d} = \frac{\mathbf{v}_{\parallel}^{2}}{\Omega}\hat{\mathbf{b}} \times (\hat{\mathbf{b}} \cdot \nabla \hat{\mathbf{b}}) + \frac{\mu}{\Omega}\hat{\mathbf{b}} \times \nabla B \times \nabla B$$

$$\mathbf{v}_{d} = \frac{\mathbf{v}_{\parallel}^{2}}{\Omega}\hat{\mathbf{b}} \times (\hat{\mathbf{b}} \cdot \nabla \mathbf{b}) + \frac{\mathbf{v}_{\parallel}^{2}}{\Omega}\hat{\mathbf{b}} \times \nabla B \times \nabla$$

details:

* this is not the angular Drift-Kinetic Eq. of
Chew, Goldberger, + Low, which was for the strong E-field
"MHO ordering" (See Kulsrad, Handbook of Plasma Physics, 1983)

VE ~ Vt >> Vd ~ VI

SIR ~ Vt R

* closer to the form of the Drift-Kinetic Eq. used in neodassical theory, where $V_E \sim V_A$ ("weak E-freld") even though $\frac{V_E}{V_t} \sim f_R \sim E$, $\frac{V_E \cdot \nabla}{V_{li} \cdot \delta \cdot \nabla} \sim \frac{V_t \cdot f_R \cdot k_L}{V_t \cdot h_{li}} \sim \frac{k_L f_R}{h_{li} \cdot K}$ ~ 1

Gyrokinetic Eq. for fill gurding-center density f(k, v11, p,t):

$$\frac{\partial \widehat{f}}{\partial t} + (v_{\parallel} \hat{\mathbf{b}} + \mathbf{v}_E + \mathbf{v}_d) \cdot \nabla \overline{f} + \left(\frac{q}{m} E_{\parallel} - \mu \nabla_{\parallel} B + v_{\parallel} (\hat{\mathbf{b}} \cdot \nabla \hat{\mathbf{b}}) \cdot \mathbf{v}_E \right) \frac{\partial \overline{f}}{\partial v_{\parallel}} = 0$$

In the uniform B slab limit, this is = to Krommer GK Eq. 4

Homework show that expanding the Boltzmann factor in

Cowley's Eq. 37, a gyro averaging to get $f = F_0 - 9 \langle \Phi \rangle F_0 + h$ above GK Eq.

$$\overline{f} = F_0 - 9 \underbrace{\langle \overline{P} \rangle}_{T_0} F_0 + h$$

gives exactly Cowley's (Frieman-Chen) form of the GK Eg. (Cowley Eq. 40) for the (Use uniform B slab limit for simplicity).

[dexpond in consistent assumptions:

$$F_0 \nabla_L \frac{g(\Phi)}{T_0} \sim \nabla_L F_0$$

$$q < \frac{1}{T} < c1$$
 but

$$F_0V_L\frac{q(F)}{T_0}\sim V_LF_0$$
 $F_0V_L\frac{q(F)}{T_0}\sim V_LF_0$

Gyrukihetic Eq. for fill gurding-center density f(R, VII, p,t):

$$\frac{\partial \overline{f}}{\partial t} + (v_{\parallel} \hat{\mathbf{b}} + \mathbf{v}_E + \mathbf{v}_d) \cdot \nabla \overline{f} + \left(\frac{q}{m} E_{\parallel} - \mu \nabla_{\parallel} B + v_{\parallel} (\hat{\mathbf{b}} \cdot \nabla \hat{\mathbf{b}}) \cdot \mathbf{v}_E \right) \frac{\partial \overline{f}}{\partial v_{\parallel}} = 0$$

Homework's Show that substituting the gyro-average of Cowley's Eq. 37:

$$\frac{1}{f} = F_0 - 9 \frac{\langle \overline{\Phi} \rangle}{T_0} F_0 + h$$

$$\frac{\partial h}{\partial t} - \frac{g}{\tau_0} \frac{\partial \Phi}{\partial t} = + v_{11} \hat{h} \cdot \nabla h + v_{E} \cdot \nabla h + v_{E} \cdot \nabla F_0 \left(1 - g \cdot \frac{\partial \Phi}{\partial t} \right)$$

There 2
$$-V_{11}\dot{b}\cdot\nabla\langle\bar{\Phi}\rangle$$
 $= 0$ $-\frac{1}{2}\dot{b}\cdot\nabla\langle\bar{\Phi}\rangle\frac{\partial h}{\partial V_{11}}$ $= 0$ $+\frac{1}{2}$ $= 0$ $= 0$ $+\frac{1}{2}$ $= 0$ $=$

Homework

$$\frac{\partial \overline{f}}{\partial t} + (v_{\parallel} \hat{\mathbf{b}} + \mathbf{v}_E + \mathbf{v}_d) \cdot \nabla \overline{f} + \left(\frac{q}{m} E_{\parallel} - \mu \nabla_{\parallel} B + v_{\parallel} (\hat{\mathbf{b}} \cdot \nabla \hat{\mathbf{b}}) \cdot \mathbf{v}_E \right) \frac{\partial \overline{f}}{\partial v_{\parallel}} = 0$$

Linearize: $\vec{f} = \vec{F}_0 + \vec{f}_1$ where \vec{F}_0 satisfies Equilibrium Eq. $\frac{\partial}{\partial L} = 0$ $\vec{E} = \vec{C}$

 $(V_{ii} \stackrel{\wedge}{b} + V_{d}) \cdot \nabla F_{b} - N \nabla_{ii} B \frac{\partial F_{b}}{\partial V_{ii}} = 0$

General Equilibrium solution could be an arbitrary function of the constants of the motion (E, P, P_{ϕ}) where $E = \frac{1}{2}mv_{11}^{2} + pB$

de Por= canonical angular momentum

But it we neglect (vol ~ f

Busically says Fo = const.

along trajectories of

banana orbits or passing

orbits in a totamati.

get simpler Eg:

$$V_{II} \hat{b} \cdot \nabla F_{0} - \mu(\hat{b} \cdot \nabla B) \frac{\partial F_{0}}{\partial V_{II}} = 0$$

Will consider Equilibrium of the form:

Exercise: Plug this in to the previous Eq. + show it is a solution.

$$\frac{\partial \widehat{f}}{\partial t} + (v_{\parallel} \hat{\mathbf{b}} + \mathbf{v}_E + \mathbf{v}_d) \cdot \nabla \overline{f} + \left(\frac{q}{m} E_{\parallel} - \mu \nabla_{\parallel} B + v_{\parallel} (\hat{\mathbf{b}} \cdot \nabla \hat{\mathbf{b}}) \cdot \mathbf{v}_E \right) \frac{\partial \overline{f}}{\partial v_{\parallel}} = 0$$

Linearize: F=Fo+F, where Fo satisfies Equilibrium Eq.

Next order Eq:

$$\frac{\partial f}{\partial f} + (v_{\parallel}b + v_{\parallel}d) \cdot \nabla f - \mu \nabla_{\parallel} B \frac{\partial f}{\partial v_{\parallel}} = - \sqrt{\epsilon} \cdot \nabla F_{0}$$

$$- (\frac{a}{m} E_{\parallel} + v_{\parallel}(b, \nabla b) \cdot v_{E}) \frac{\partial F_{0}}{\partial v_{\parallel}}$$

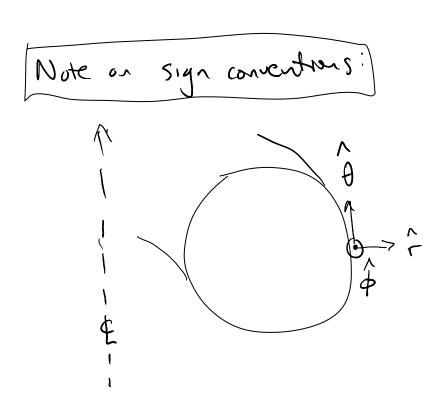
$$- (\frac{b}{m} E_{\parallel} + v_{\parallel}(b, \nabla b) \cdot v_{E}) \frac{\partial F_{0}}{\partial v_{\parallel}}$$

$$\left(- \lambda \omega + i \nu_{\parallel} h_{\parallel} + i \nu_{d} \cdot h_{\perp} \right) \hat{f} = - \nu_{E} \cdot \nabla F_{o}$$

$$- \left(\frac{a}{m} E_{\parallel} + \nu_{\parallel} (b \cdot \nabla b) \cdot \nu_{E} \right) \frac{\partial F_{o}}{\partial \nu_{\parallel}}$$

Important Subtlety:
$$F(R, V_n, p, t)$$
 so

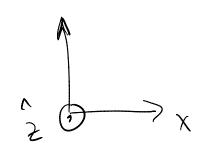
 $-V_E \cdot \nabla F_0 = -V_E \cdot \nabla | F_0$
 $V_{n,p,t}$
 V_{n,p



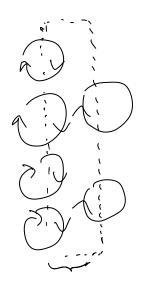
defining $\omega_{dv} = h \cdot Vd$ gives convention used in Beer's thesis!

$$\omega_{dv} = \omega_d (v_{\parallel}^2 + \mu B) / v_t^2$$
$$\omega_d = -k_{\theta} \rho v_t / R$$

More on Sign Conventions



E Vn



with B out of page, the diamagnetic flow Vxi is downward if Rn is inward. Thus

$$\omega_{\star i} = h \cdot V_{\star i} = -h_{\theta} V_{t} f_{h}$$

$$= - \frac{cT}{eB} \frac{h_0}{L_r}$$

(Back to RHS of Incarized GK Eq., 4 slides back)

$$RHS = -V_{E} \cdot \nabla F_{o} - \left(\frac{2}{m} E_{1} + V_{11}(\hat{b} \cdot \nabla \hat{b}) \cdot V_{E}\right) \frac{2F_{o}}{2V_{11}}$$

$$\sim -\frac{c}{B} \nabla F \times \hat{b} \cdot \rho \nabla B$$

$$\sim -\nabla F \cdot \left[\rho \cdot \nabla \hat{b} \times \nabla B + V_{11}^{2} \cdot \left(\hat{b} \cdot \nabla \hat{b}\right) \cdot \left(\frac{\hat{b} \times \nabla F}{B}\right)\right]$$

$$\sim -\nabla F \cdot \left[\rho \cdot \nabla \hat{b} \times \nabla B + V_{11}^{2} \cdot \left(\hat{b} \cdot \nabla \hat{b}\right) \cdot \left(\frac{\hat{b} \times \nabla F}{B}\right)\right]$$

$$\sim -\nabla F \cdot \left[\rho \cdot \nabla \hat{b} \times \nabla B + V_{11}^{2} \cdot \left(\hat{b} \cdot \nabla \hat{b}\right) \cdot \left(\frac{\hat{b} \times \nabla F}{B}\right)\right]$$

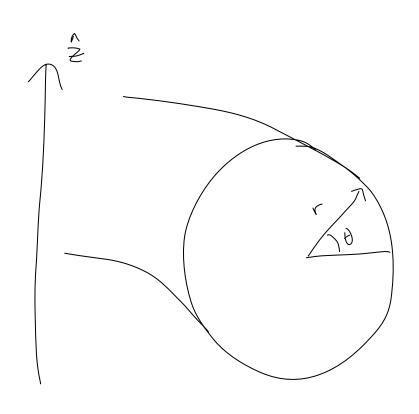
RHS =
$$+i\left(\omega_{xv}^{T} - \omega_{dv} - h_{ii}v_{ii}\right)\frac{e\overline{F}}{T_{o}}F_{o}$$

$$\omega_*^T = \omega_* [1 + \eta (v_{\parallel}^2 / 2v_t^2 + \mu B / v_t^2 - 3/2)]$$

$$\omega_{*} = h_{\theta} p \frac{V_{t}}{L_{n}}$$
 $\eta = \frac{L_{n}}{L_{\tau}}$

$$\omega_{dv} = \omega_d(v_{\parallel}^2 + \mu B)/v_t^2$$

$$\omega_{a} = -\frac{v_{t}}{R} \rho \left(h_{\theta} \cos \theta + h_{r} \sin \theta \right)$$



downward

Vd from VB + cunature

drift

 $\omega_{\mathbf{d}} = \frac{h \cdot Vd}{L}$ $= -\frac{V+L}{R} \left(h_{\theta} \cos \theta + h_{r} \sin \theta \right)$

will focus on 0 = 0 here (where bad-curvature drive is the strongest)

$$\left(-\lambda \omega + i V_{11} h_{11} + i V_{2} \cdot h_{11} \right) \tilde{f} = -V_{E} \cdot \nabla F_{o}$$

$$-\left(\frac{2}{m} E_{11} + V_{11} \left(h_{2} \nabla h_{2} \right) \cdot V_{E} \right) \frac{\partial F_{o}}{\partial V_{11}}$$

$$\left(-\lambda \omega + i V_{11} h_{11} + i \omega_{dV} \right) \tilde{f} = -i \left(-\omega_{*v}^{T} + \omega_{dv} + h_{11} V_{11} \right) \frac{e \overline{F}}{T_{o}} F_{o}$$

$$\tilde{f} = \frac{-\omega_{*v}^{T} + \left(h_{11} V_{11} + \omega_{dv} \right)}{\omega - \left(h_{11} V_{11} + \omega_{dv} \right)} \frac{e \overline{\Phi}}{T_{o}} F_{o}$$

Note: recover Boltzmann response when hy VII dor Wou large

$$\frac{\partial}{\partial t} = \frac{-\omega_{xv}^T + (h_{11}v_{11} + \omega_{dv})}{\omega - (h_{11}v_{11} + \omega_{dv})} = \frac{\partial}{\partial t} F_0$$

Look for modes with

hyVti << W, Wxv, Wav

< c h, Vte

(slab "n;" version of ITG
requires finite him toi, but not
toroidal version),

assume Boltzmann electrons

Quasinestrality:

(additional polarization contribution to density gues his corrections but not critical for basic ITE.)

$$N_{eo} = \int d^3 \frac{-\omega_{\star v} + \omega_{av}}{\omega - \omega_{dv}} F_o = \int \frac{e^{\frac{\pi}{2}}}{T_{av}}$$

$$N_{o} = \frac{1}{T_{eo}} = N_{o} = \frac{1}{T_{o}} \int_{0}^{3} \frac{V_{dv} - \omega_{xr}}{\omega_{o}} = \frac{\omega_{dv} - \omega_{xr}}{\omega_{o}}$$

"Cold plasma" or "fast wave" approx. W>> Wav

$$\frac{T_{no}}{T_{eo}} = \int \int_{N_0}^{3} \frac{F_o}{\omega} \frac{\omega_{dv} - \omega_{xt}}{\omega} \left(1 + \frac{\omega_{dv}}{\omega} + \cdots \right)$$

$$\frac{T_{\text{No}}}{T_{\text{eo}}} = \int \int_{\text{No}}^{3} \frac{F_{\text{o}}}{\kappa_{\text{o}}} \frac{\omega_{\text{dV}} - \omega_{\text{XT}}}{\omega} \left(1 + \frac{\omega_{\text{dV}}}{\omega} + \cdots \right)$$

$$\omega_{\text{dv}} = \omega_{\text{d}}(v_{\parallel}^{2} + \mu B)/v_{t}^{2} \qquad \omega_{*}^{T} = \omega_{*}[1 + \eta(v_{\parallel}^{2}/2v_{t}^{2} + \mu B/v_{t}^{2} - 3/2)]$$

$$\omega_{\text{d}} = -k_{\theta}\rho v_{t}/R \qquad \omega_{*} = -k_{\theta}\rho v_{t}/L_{n}$$

$$= v_{\text{X}}^{2} + v_{\text{Y}}^{2}$$

$$(13) F_{\text{o}} = (13) F_{\text{o}} = (13$$

$$\int d^3 v \frac{F_0}{N_0} \omega_{dv} = \int d^3 v \frac{F_0}{N_0} \omega_d \left(v_{11}^2 + \frac{1}{2} v_{\perp}^2 \right) / v_t^2$$

$$= 2 \omega_d$$

Using useful I.D. for Maxwellian Fo:
$$\langle V_{x}^{2n} \rangle = \int d^{3}y \frac{F_{o}}{n_{o}} V_{x}^{2n} = V_{t}^{2n} \frac{(2n-1)!!}{(2n-3)(2n-5)\cdots 5\cdot 3\cdot 1}$$

$$\frac{T_{00}}{T_{00}} = \int d^{3}v \frac{F_{0}}{N_{0}} \frac{\omega_{0}v - \omega_{*T}}{\omega} \left(1 + \frac{\omega_{0}v}{\omega} + \cdots \right)$$

$$\omega_{dv} = \omega_{d}(v_{\parallel}^{2} + \mu B)/v_{t}^{2} \qquad \omega_{*}^{T} = \omega_{*}[1 + \eta(v_{\parallel}^{2}/2v_{t}^{2} + \mu B/v_{t}^{2} - 3/2)]$$

$$\omega_{d} = -k_{\theta}\rho v_{t}/R \qquad \omega_{*} = -k_{\theta}\rho v_{t}/L_{n} \qquad \qquad = \frac{1}{2}v_{\perp}^{2} = \frac{1}{2}\left(v_{\chi}^{2} + v_{y}^{2}\right)$$

$$\int d^{3}v \frac{F_{0}}{N_{0}} \omega_{*}^{T} = \omega_{*}\left(1 + \eta\left(\frac{1}{2} + 1 - \frac{3}{2}\right)\right) = \omega_{*}$$

$$\int d^{3}v \frac{F_{0}}{N_{0}} \omega_{dv}^{T} = \int d^{3}v \frac{F_{0}}{N_{0}} \omega_{d}^{2}\left[v_{u}^{T} + 2v_{u}^{2} + v_{x}^{2}v_{x}^{2} + v_{y}^{2}\right)^{2}\right] \frac{1}{V_{t}^{2}}$$

$$= \omega_{d}^{2}\left[3 + 2 \cdot \frac{1}{2}\left(1 + 1\right) + \frac{1}{4}\left(\langle v_{x}^{2} + v_{x}^{2} \rangle_{v}^{2} + v_{y}^{4}\rangle\right)\right]$$

$$= \omega_{d}^{2}\left[5 + \frac{1}{4}\left(8\right)\right] = 7\omega_{d}^{2}$$

$$\frac{T_{no}}{T_{eo}} = \int \int_{N_0}^{3} \frac{F_o}{\omega} \frac{\omega_{dv} - \omega_{*T}}{\omega} \left(1 + \frac{\omega_{dv}}{\omega} + \cdots \right)$$

$$\omega_{dv} = \omega_d(v_{\parallel}^2 + \mu B)/v_t^2 \qquad \omega_*^T = \omega_* [1 + \eta(v_{\parallel}^2/2v_t^2 + \mu B/v_t^2 - 3/2)]$$

$$\omega_d = -k_{\theta}\rho v_t/R \qquad \omega_* = -k_{\theta}\rho v_t/L_n \qquad \qquad = \frac{1}{2} v_{\perp}^2 = \frac{1}{2} \left(v_{\parallel}^2 + v_{\parallel}^2\right)$$

$$\int d^{3}v \frac{F_{o}}{N_{o}} \omega_{dv} \omega_{*}^{T} = \omega_{d}\omega_{*} \left\{ 2 + \frac{1}{2}v_{1}^{2} \right\} \left(\frac{1}{2}v_{11}^{2} + \frac{1}{2}v_{1}^{2} - \frac{3}{2}v_{t}^{2} \right)$$

$$+ \eta \int d^{3}v \frac{F_{o}}{N_{o}} \left(\frac{v_{11}^{2} + \frac{1}{2}v_{1}^{2}}{v_{t}^{2}} \right) \left(\frac{1}{2}v_{11}^{2} + \frac{1}{2}v_{1}^{2} - \frac{3}{2}v_{t}^{2} \right)$$

$$= \omega_{d}\omega_{*} \left\{ 2 + \eta \left[\frac{1}{2} 3 + \frac{1}{2} 2 - \frac{3}{2} + \frac{1}{2} \cdot 2 \cdot \frac{1}{2} + \frac{1}{4} 8 \right] - \frac{1}{2} \cdot 2 \cdot \frac{3}{2} \right\}$$

$$\int d^{3}v \frac{F_{o}}{N_{o}} \omega_{dv} \omega_{\star}^{T}$$

$$= \omega_{d} \omega_{\star} \left\{ 2 + \eta \left[\frac{1}{2} 3 + \frac{1}{2} 2 - \frac{3}{2} + \frac{1}{2} \cdot 2 \cdot \frac{1}{2} + \frac{1}{4} 8 \right] \right\}$$

$$= \omega_d \omega_* 2 (lt \eta)$$

Combine results from last 2 pages:

$$\frac{T_{10}}{T_{eo}} = 2 \frac{\omega_d}{\omega} - \frac{\omega_*}{\omega} + 7 \frac{\omega_d}{\omega^2} - 2 \frac{\omega_d \omega_*}{\omega^2} (1+\eta)$$

This defines a dispersion relation w us. h

$$\frac{T_{10}}{T_{eo}} = 2 \frac{\omega d}{\omega} - \frac{\omega_{\star}}{\omega} + 7 \frac{\omega_{d}}{\omega^{2}} - 2 \frac{\omega_{d} \omega_{\star}}{\omega^{2}} (1+\eta)$$

Consider the flat density limit: $\nabla n \rightarrow 0$, but $\nabla T \neq 0$ $\omega_{*} = -h_{\theta} \rho \frac{V_{t}}{L_{n}} \rightarrow 0$ $\eta = \frac{1}{L_{n}} \nabla r = \frac{L_{n}}{L_{n}} \rightarrow \infty$

 $\omega_{\star}\eta = -h_{\theta}\rho \frac{V_{t}}{L_{n}} \frac{L_{n}}{L_{\tau}} \equiv \overline{\omega}_{\star\tau}$

 $\omega^2 \frac{T_{io}}{T_{eo}} = 2 \omega_d \omega + 2 \omega_d \overline{\omega}_{x\tau} - 7 \omega_d^2 = 0$

$$W = 2 \omega_0 \pm \sqrt{4 \omega_a^2 - 4 \frac{T_{io}}{T_{e_o}} (2 \omega_a \overline{\omega_{\star t}} - 7 \omega_a^2)}$$

$$2(T_{io}/T_{e_o})$$

From last page:

$$W = 2 \omega_0 \pm \sqrt{4 \omega_a^2 - 4 \frac{T_{iv}}{T_{eo}} (2 \omega_a \omega_{\star\tau} - 7 \omega_a^2)}$$

$$2 \left(\frac{T_{iv}}{T_{eo}} \right)$$

Consider large temperature gradient limit: WxT ~ VT 1 Growth rate:

Y =
$$\frac{\sqrt{2} \omega_d \overline{\omega}_{*T}}{\sqrt{T_{AO}/T_{eO}}} = \frac{\sqrt{2} h_{\theta} p_i}{\sqrt{T_{AO}/T_{eO}}} \frac{V_{t,i}}{\sqrt{RL_T}}$$

Fundamental scaling of bad-curvature driven instabilities. Go back to general D.R.:

$$\omega = 2 \omega_0 \pm \sqrt{4 \omega_a^2 - 4 \frac{\text{Tio}}{\text{Te}_o} (2 \omega_a \overline{\omega}_{*T} - 7 \omega_a^2)}$$

$$2(\text{Tio}/\text{Te}_o)$$

$$= 2 \omega_{d} \pm \sqrt{\left(4 + 28 \frac{T_{no}}{T_{eo}}\right) \omega_{d}^{2} - 8 \frac{\overline{T_{no}}}{T_{eo}} \omega_{d} \overline{\omega}_{\star T}}$$

2 (T, 1/Teu)

Instability exists if

$$\frac{1}{R} \frac{1}{L_{T}} \Rightarrow \frac{1}{R^{2}} \left(\frac{1}{2} \frac{T_{eo}}{T_{ro}} + \frac{1}{2} 7 \right)$$

$$\frac{R}{L_{T}} \Rightarrow \frac{1}{2} \left(7 + \frac{T_{eo}}{T_{ro}} \right)$$

Note: (1) To reduce growth rate far above marginal stability, want to reduce omega_d ~ 1/R, but (2) to raise the instability threshhold, want to raise omega_d ~ 1/R



Compare w/Romanelli 1990 (tq.12):
$$\eta_i = (\tfrac{5}{3} + \tau/4) 2\epsilon_n$$

$$\frac{L_{1}}{L_{4}} = \left(\frac{5}{3} + \frac{1}{4} + \frac{T_{e}}{T_{i}}\right)^{2} \frac{L_{n}}{R}$$

$$= 3.33 + 0.5 \frac{\text{Te}}{\text{Tio}}$$

Note there is an instability only if $\omega_a \overline{\omega}_{x\tau} > 0$ $\omega_a \overline{\omega}_{*7} = (h_{\theta} \rho)^2 \frac{\overline{V_t}}{R L_T}$ unstable "bad curvature Side" Stuble curvature
good side

the Tio dependence of Teo Roe accurate: R > R = Because near marginal stubility, the of the resonant denominator $\frac{1}{\omega} \approx \frac{1}{\omega} \left(1 + \frac{\omega_{dv}}{\omega} + \cdots \right)$ w ~ Wd new Morgmal stability... breaks down, since

To get this more accurately, need to include resonance effects. Can write the exact plasma response in terms of the Z function, without expanding 1/(omega-omega_dv), see Beer and Hammett 1996, "Toroidal gyrofluid equations for simulations of tokamak turbulence", Phys. Plasmas 3, 4046, and references therein. This introduces stabilizing effects from Landau damping from the spread in drift velocities in omega_dv, which increase with T_i, causing the critical R/L_Tcrit to increase at higher T_i.

More general result for threshold for mstability! $\frac{R_o}{L_{Tcnit}} = M_{ax} \left[\left(1 + \frac{T_i}{T_e} \right) \left(1.33 + 1.91 \frac{S}{q} \right) \left(1.5 \frac{r}{R} \right) \left(1+0.3 \frac{r d K}{d r} \right) \right]$ $0.8 \frac{R_o}{L_h}$

Found by fits to lots of GS2 Gyrokinetic stability calculations (Jenko, Dorland Hommett, PoP 2001), guided by previous analytic results (Romanelli, Hahm + Tang) in some limits.

ITG References

- Mike Beer's Thesis 1995 http://w3.pppl.gov/~hammett/collaborators/mbeer/afs/thesis.html
- Romanelli & Briguglio, Phys. Fluids B 1990
- Biglari, Diamond, Rosenbluth, Phys. Fluids B 1989
- Jenko, Dorland, Hammett, PoP 2001
- Candy & Waltz, PRL ...
- Kotschenreuther et al.
- Dorland et al, PRL ...
- Dimits et al....
- ...
- Earlier history:
 - slab eta_i mode: Rudakov and Sagdeev, 1961
 - Sheared-slab eta_i mode: Coppi, Rosenbluth, and Sagdeev, Phys. Fluids 1967
 - Toroidal ITG mode: Coppi and Pegoraro 1977, Horton, Choi, Tang 1981, Terry et al. 1982, Guzdar et al. 1983... (See Beer's thesis)